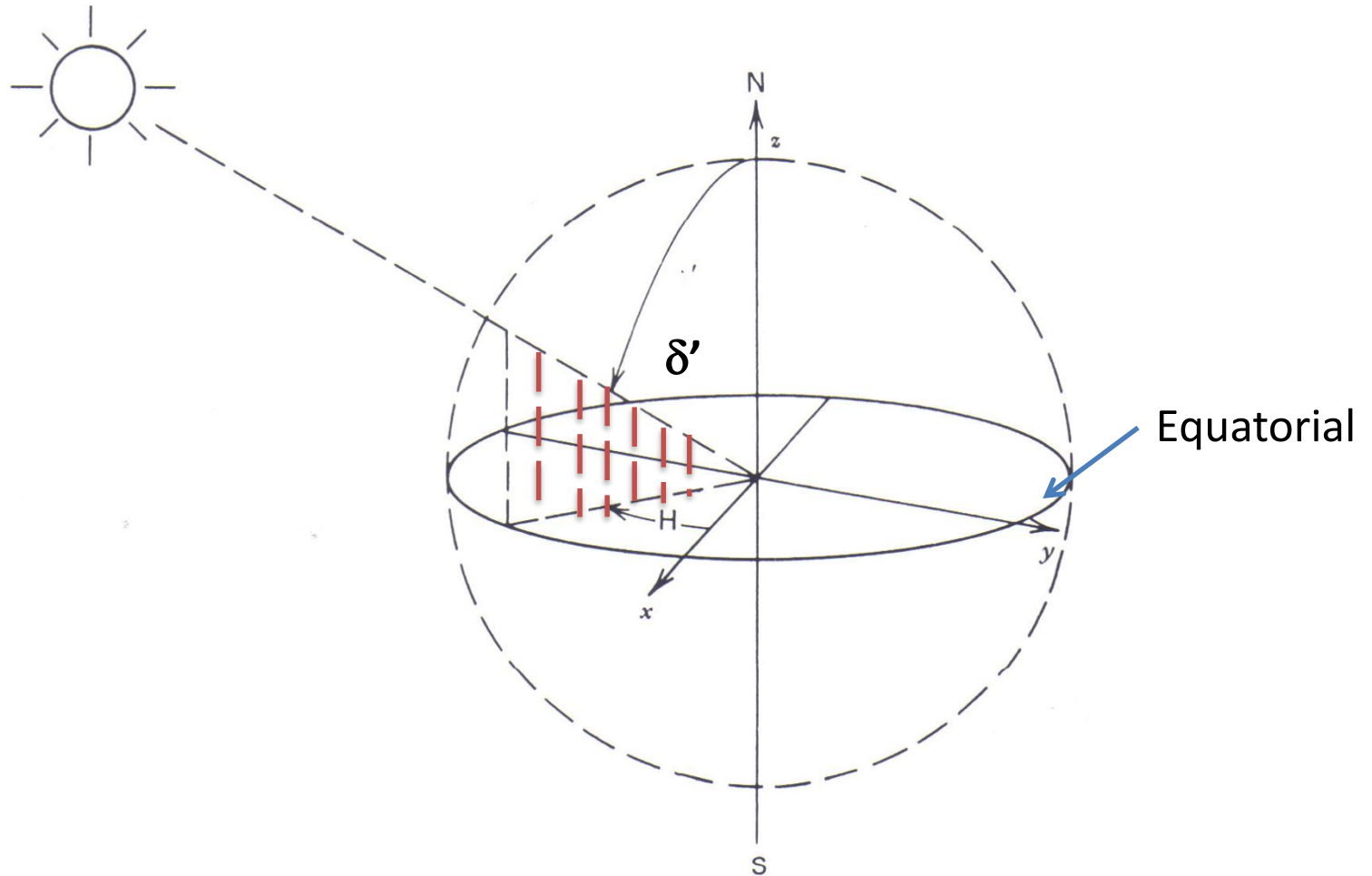


Solar Coordinates & Solar Insolation

Solar Coordinates

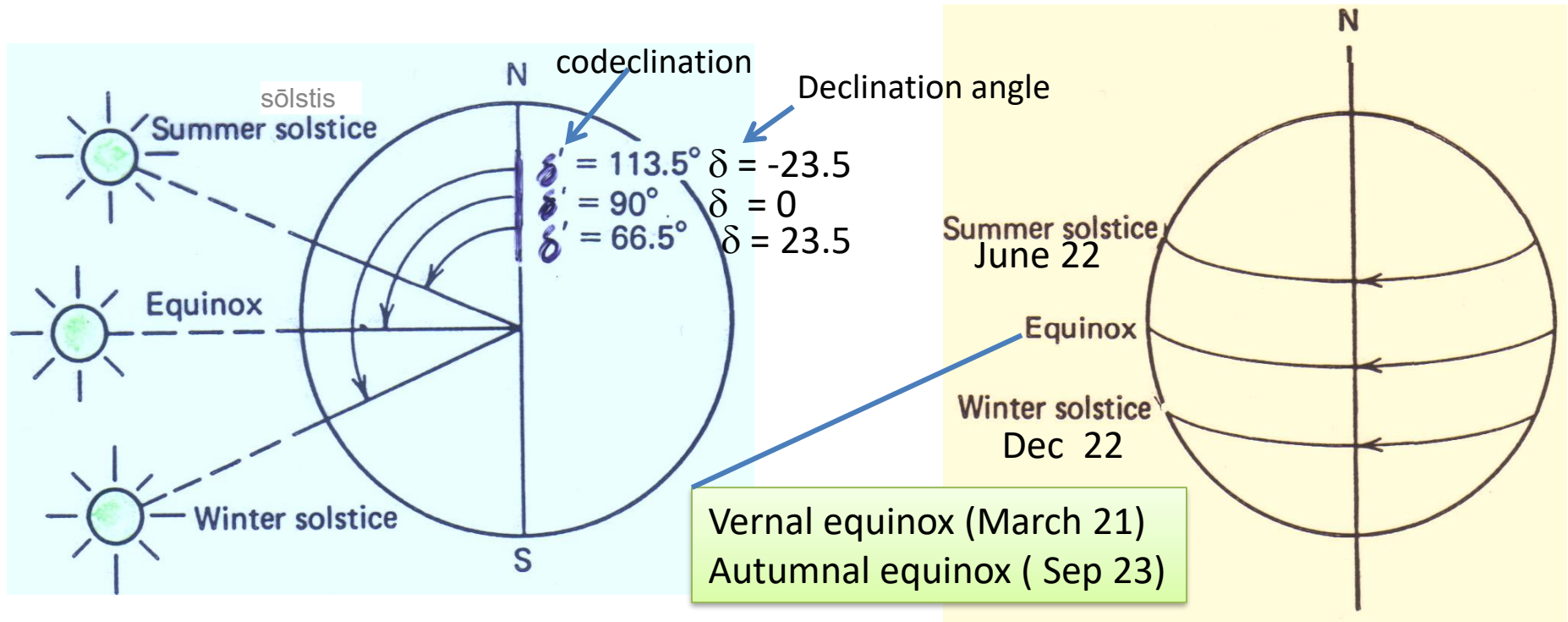
- Geocentric coordinate system based on earth's center!
- The position of the sun can be ultimately determined by two angles:
 1. codeclination δ' (complement of declination δ).
 δ' varies from 66.5° at the summer solstice (June 22) to 113.5° at the winter solstice (Dec. 22).
 2. hour angle H , $= \pm 360t/24\text{hr}$ where t is the number of hours before or after solar noon. –ve refers to time before solar noon.

Solar Coordinates



Geocentric coordinate system

The seasonal variation of the codeclination.



The seasonal variation of the angle between the earth's polar axis and the earth-sun line.

Note

$$\cos \delta' = \sin 23.5^\circ \sin \frac{360^\circ n}{365.25 \text{ days}}$$

$$\delta = 90^\circ - \delta'$$

Where n is the number of days after the vernal equinox (March 21).

Also δ can be found from tables (solar tables)

Local Solar Coordinates

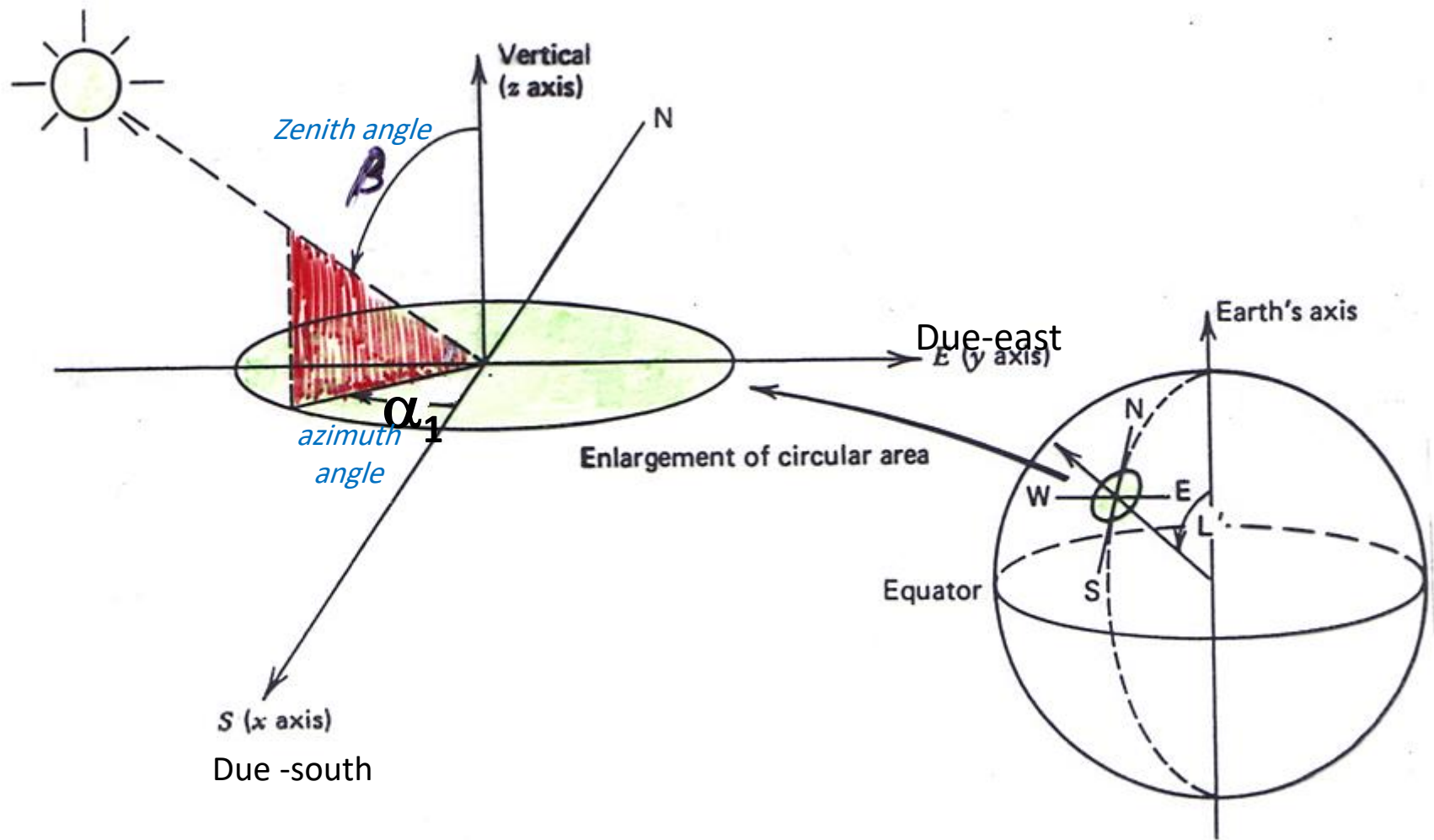
- Geocentric sys. is not suitable for the observer on the earth surface.
- The following coordinate is more appropriate:
 1. zenith angle β and its complement is called the altitude of the sun β_1
 2. azimuth angle α_1 .

Due – south \sim x axis

Due –east \sim y axis

Vertical to equator is z axis

Local solar coordinates—the zenith angle β and the azimuth angle α_1



Altitude (β_1) and Azimuth (α_1) angles “Solar angles”

$$\sin \beta_1 = (\cos L)(\cos \delta)(\cos H) + (\sin L)(\sin \delta)$$

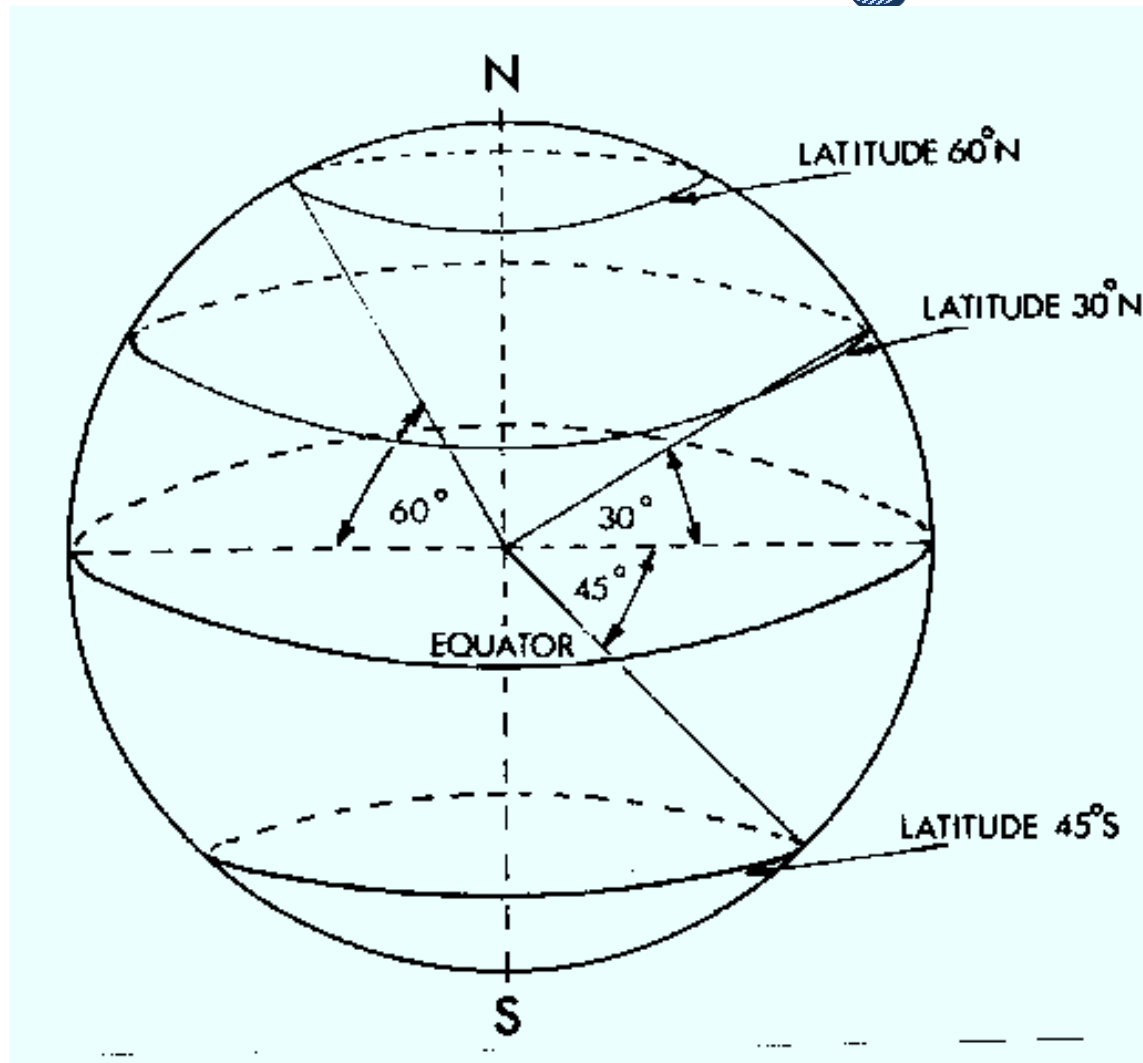
$$\sin \alpha_1 = \frac{\cos \delta \sin H}{\cos \beta_1}$$

Where L is the Latitude angle

Note

The values of α_1 and β_1 are given in the recommended text.

Latitude Angle

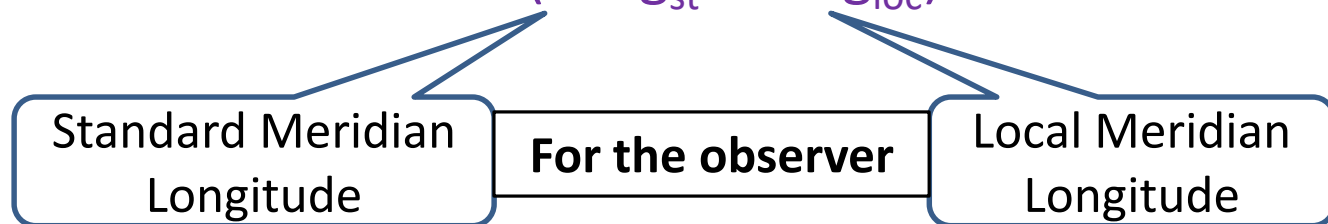


Solar time

- Equation of time (EOT): it is the cumulative variation of the solar noon. The following table shows the EOT in minutes.

Day of Month	January	February	March	April	May	June	July	August	September	October	November	December
1	-4	-14	-13	-4	+3	+2	-3	-6	0	+10	+16	+11
4	-5	-14	-12	-3	+3	+2	-4	-6	+1	+11	+16	+10
7	-6	-14	-11	-2	+3	+2	-5	-6	+2	+12	+16	+9
10	-8	-14	-10	-1	+4	+1	-5	-5	+3	+13	+16	+7
13	-9	-14	-10	-1	+4	0	-6	-5	+4	+14	+16	+6
16	-10	-14	-9	0	+4	0	-6	-4	+5	+14	+15	+4
19	-11	-14	-8	+1	+4	-1	-6	-4	+6	+15	+15	+3
22	-12	-14	-7	+1	+4	-2	-6	-3	+7	+15	+14	+2
25	-12	-13	-6	+2	+3	-2	-6	-2	+8	+16	+13	0
28	-13	-13	-5	+2	+3	-3	-6	-1	+9	+16	+12	-2

- Solar Noon = $12:00 - 0:04(\text{Long}_{\text{st}} - \text{long}_{\text{loc}}) - \text{EOT}$



- Mean Sun Time, $MST = \text{local standard time} + 0:04(\text{Long}_{st} - \text{long}_{loc})$
- Apparent Sun Time, $AST = MST + EOT$

Important for hour
angle

- Examples : Standard meridian for USA
 - 75 °W Long. (EST), Eastern standard meridian
 - 90 ° W Long.(CST), Central standard meridian
 - 105 ° W Long. (MST), mountain standard meridian
 - 120 ° W Long. (PST), Pacific standard meridian

Parameters for solar calculations (on the 21st day of each month)

Month	January	February	March	April	May	June	July	August	September	October	November	December
Day of the year	21	52	80	111	141	173	202	233	265	294	325	355
Declination, degrees	-19.9	-10.6	0.0	+11.9	+20.3	+23.45	+20.5	+12.1	0.0	-10.7	-19.9	-23.45
Equation of time, min	-11.2	-13.9	-7.5	+1.1	+3.3	-1.4	-6.2	-2.4	+7.5	+15.4	+13.8	+1.6
Solar noon	Late			Early		Late			Early			
A , Btu/h·ft ² †	390	385	376	360	350	345	344	351	365	378	387	391
B , 1/m	0.142	0.144	0.156	0.180	0.196	0.205	0.201	0.177	0.160	0.149	0.149	0.142
C , dimensionless	0.058	0.060	0.071	0.097	0.121	0.134	0.136	0.122	0.092	0.073	0.063	0.057

† A is the apparent solar irradiation at air mass zero for each month.

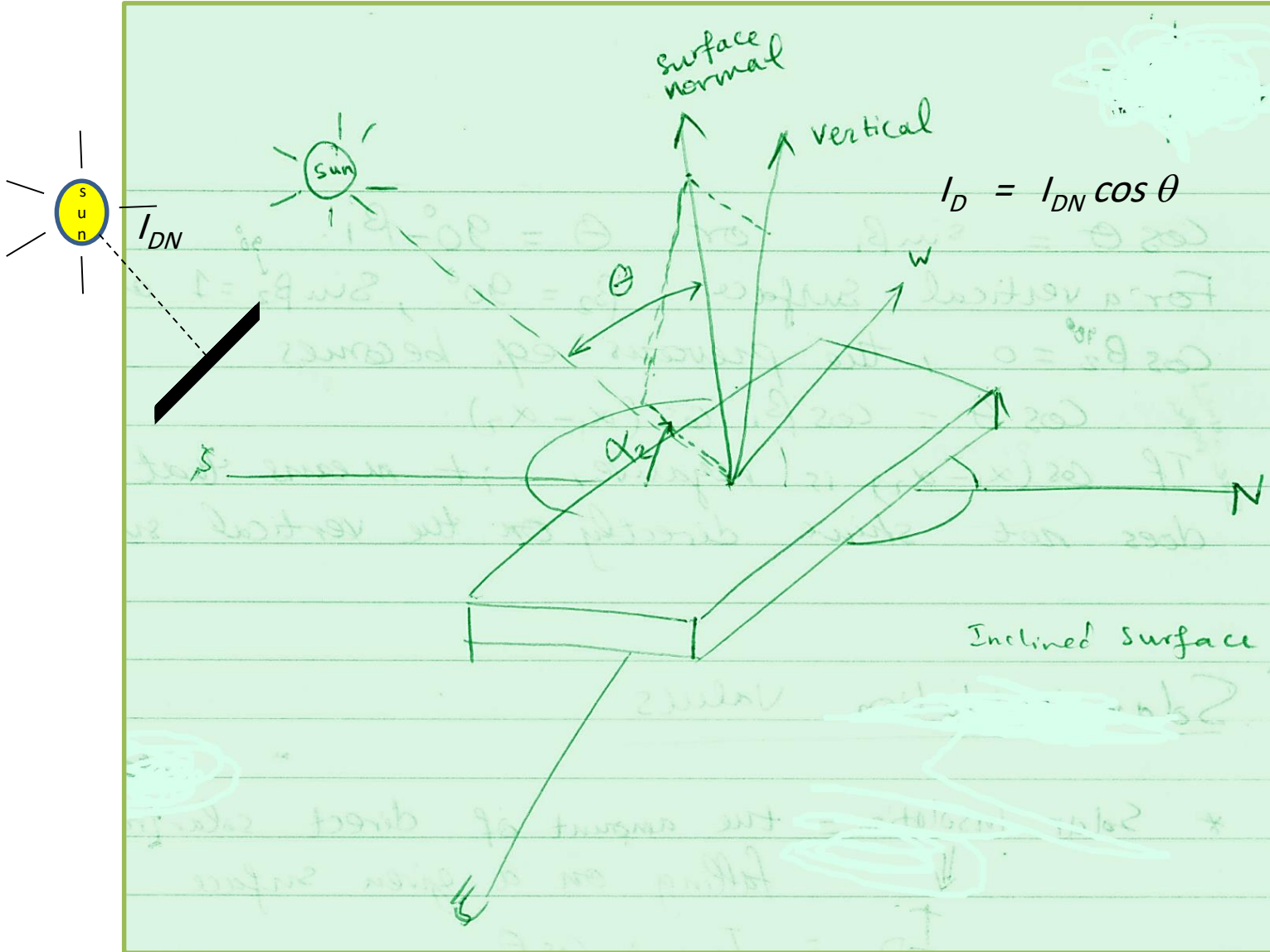
B is the atmospheric extinction coefficient.

C is the ratio of the diffuse to direct normal irradiation on a horizontal surface.

Solar Insolation “ I_D ”

Definition: **Solar Insolation** (incident or incoming **solar radiation**) is the total amount of solar radiation energy received on a given surface area during a given time. It is also called **solar irradiation** and expressed as "hourly irradiation" if recorded during an hour or "daily irradiation" if recorded during a day

Solar Insolation " I_D "_Inclined surface



Determination of the angle between the sun's rays and the normal to the surface " θ "

Firstly the orientation of the surface must be specified:

- I. α_2 : the azimuth angle of the surface. It is the angle between the horizontal projection of the normal to the surface and the due-south line measured in a clockwise direction. (See the previous slid)
- II. β_2 : tilt angle of the surface. It is the angle between the surface and the horizontal.

θ can be determined from the following equation:

$$\cos \theta = \sin \beta_1 \cos \beta_2 + \cos \beta_1 \sin \beta_2 \cos (\alpha_1 - \alpha_2)$$

Summary
Solar Insolation & Collector
Energy Absorption

Solar Insolation Values

- Total solar energy flux, I_t

$$I_t = I_D + I_{DS} + I_R \quad (1)$$

- Where I_D is the direct solar beam

I_{DS} is the diffuse component

I_R is the reflected solar radiation from surroundings.

Note: The third component, I_R , is found from the laws of heat transfer by radiation.

Direct solar energy flux

- In general, the direct solar energy flux falling on a given surface is equal to the product of the direct radiation falling on a surface normal to the sun's rays, I_{DN} , and $\cos \theta$ or:

$$I_D = I_{DN} \cos \theta$$

- I_{DN} can be estimated from:

$$I_{DN} = A_s \exp \frac{-B_s}{\sin \beta_1}$$

- Where

A_s is the apparent extraterrestrial insolation.

B_s is the atmospheric extinction coefficient. It depends on the time of year and the amount of water vapor present in the average atmosphere.

- Both of A_s and B_s are given in tables (solar tables). See next slides.

Date	Angle of declination δ	Equation of time, min	$A_S,^\dagger$ $\text{kW/m}^2\ddagger$	$B_S,^\dagger$ am§	C_S^\dagger	Day of year
January						
1	-23.0	-3.6	1231	0.142	0.057	1
11	-21.7	-8.0	1230	0.142	0.058	11
21	-19.6	-11.4	1229	0.142	0.058	21
31	-17.3	-13.5	1224	0.142	0.058	31
February						
1	-17.0	-13.6	1223	0.143	0.059	32
11	-13.9	-14.4	1218	0.143	0.059	42
21	-10.4	-13.8	1213	0.144	0.060	52
28	-7.6	-12.6	1206	0.147	0.063	59
March						
1	-7.4	-12.5	1205	0.148	0.063	60
11	-3.5	-10.2	1195	0.152	0.067	70
21	0.0	-7.4	1185	0.156	0.071	80
31	4.3	-4.3	1169	0.164	0.079	90
April						
1	4.7	-4.0	1167	0.165	0.080	91
11	8.5	-1.1	1151	0.172	0.088	101
21	12.0	1.2	1135	0.180	0.096	111
30	14.9	2.8	1126	0.185	0.104	120

Date	Angle of declination δ	Equation of time, min	A_s, \dagger $\text{kW/m}^2 \ddagger$	B_s, \dagger am§	$C_s \dagger$	Day of year
May						121
1	15.2	2.9	1125	0.186	0.105	131
11	17.9	3.7	1114	0.191	0.113	141
21	20.3	3.6	1103	0.196	0.121	151
31	22.0	2.5	1098	0.199	0.125	
June						152
1	22.1	2.4	1097	0.199	0.126	162
11	23.1	0.6	1092	0.202	0.130	172
21	23.45	-1.5	1087	0.205	0.134	181
30	23.1	-3.4	1086	0.206	0.135	
July						182
1	23.1	-3.4	1086	0.206	0.135	192
11	22.1	-5.6	1085	0.206	0.135	202
21	20.6	-6.2	1084	0.207	0.136	212
31	18.1	-6.2	1091	0.205	0.132	
August						213
1	17.9	-6.2	1092	0.205	0.132	223
11	15.2	-5.1	1099	0.203	0.126	233
21	12.0	-3.1	1106	0.201	0.122	243
31	8.7	-1.5	1120	0.193	0.113	
September						244
1	8.2	0.0	1122	0.192	0.112	254
11	4.4	3.3	1136	0.185	0.097	264
21	0.0	6.8	1150	0.177	0.092	273
30	-2.9	9.9	1162	0.172	0.086	
October						274
1	-3.3	10.2	1164	0.171	0.086	284
11	-7.2	13.1	1177	0.166	0.079	294
21	-10.8	15.3	1191	0.160	0.073	304
31	-14.3	16.2	1200	0.157	0.070	
November						305
1	-14.6	16.3	1201	0.156	0.069	315
11	-17.5	15.9	1211	0.153	0.066	325
21	-20.0	15.1	1220	0.149	0.063	334
30	-21.7	11.3	1224	0.147	0.061	
December						335
1	-21.9	11.0	1224	0.147	0.061	345
11	-23.0	6.8	1228	0.144	0.059	355
21	-23.45	2.0	1232	0.142	0.057	365
31	-23.1	-3.1	1231	0.142	0.057	

(\dagger : atmosphere)

Diffuse component, I_{DS}

- The diffuse component can be estimated from

$$I_{DS} = C_s I_{DN} F_{ss}$$

Where C_s is the ratio of diffuse to direct solar radiation falling on a horizontal surface. Its value is given in the previous tables.

F_{ss} is the radiation shape factor for radiation leaving the surface that hits the sky.

F_{ss} can be approximated by

$$F_{ss} = \frac{1 + \cos \beta_2}{2}$$

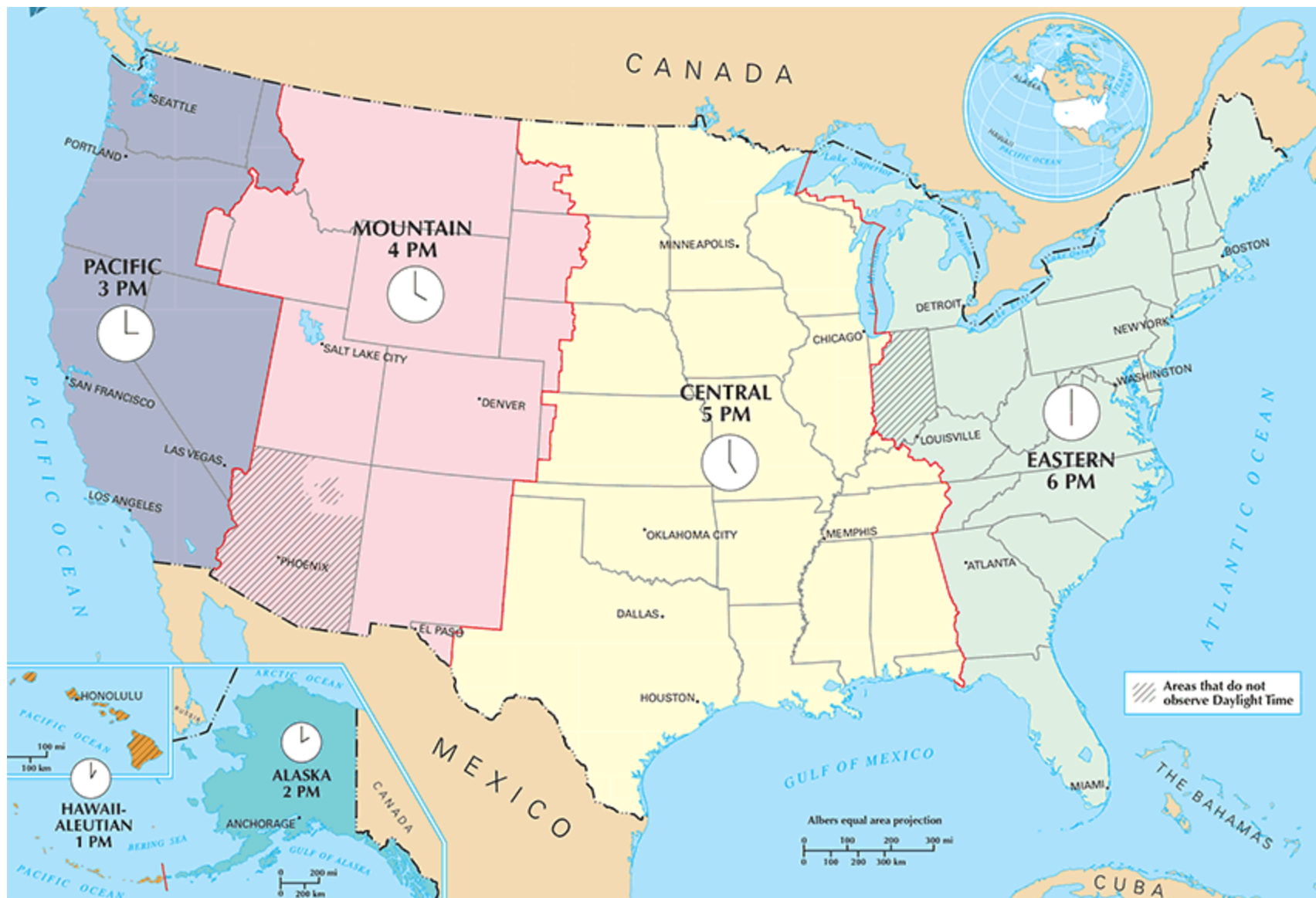
Solar energy absorbed by a given surface

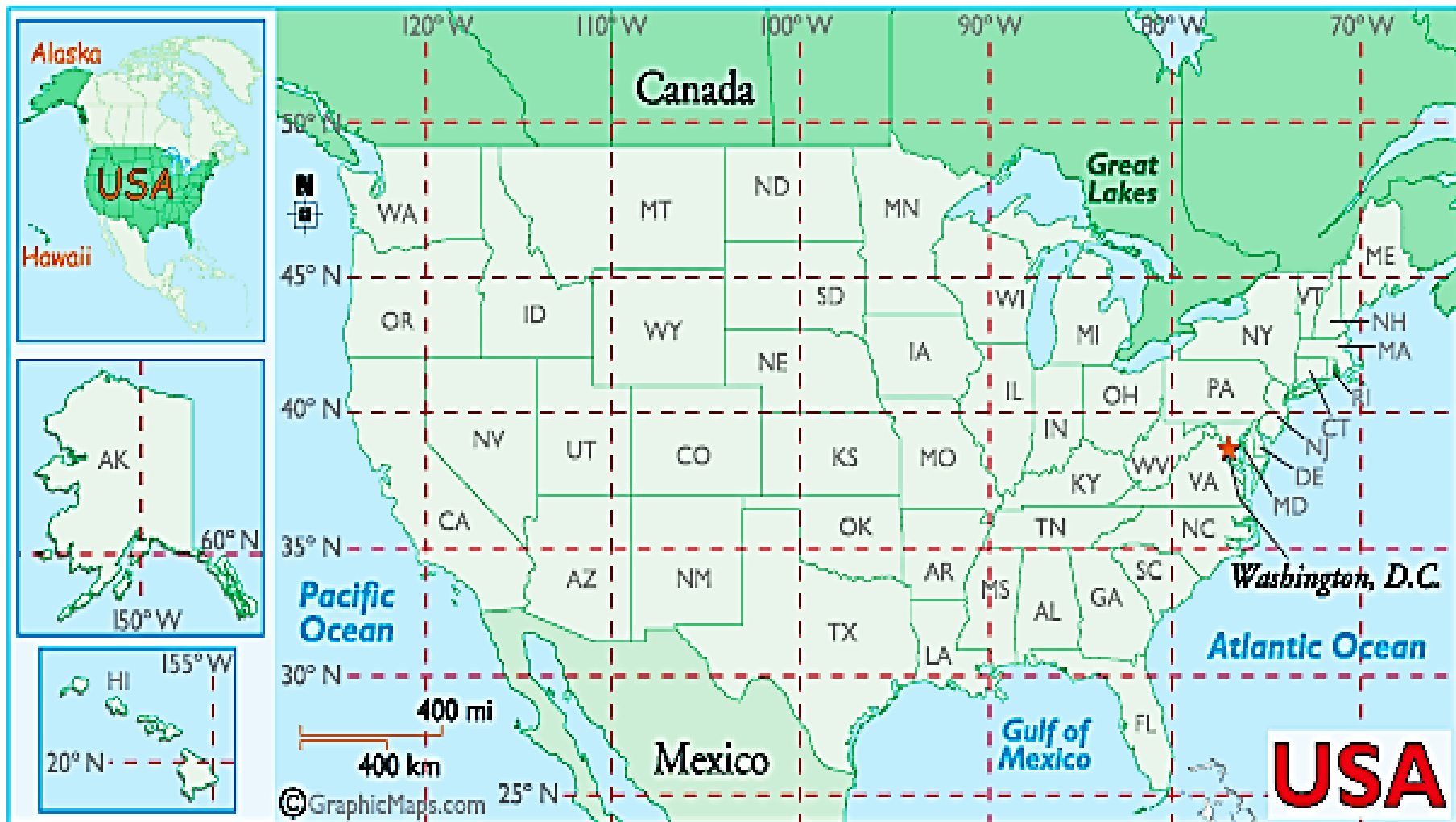
- *The product of I_t and the absorptivity or emissivity of the surface or:*

$$\text{Absorbed solar energy flux} = I_t * \varepsilon$$

Solar Emissivities

Solar emissivities	Infrared (low-temp.) radiation	Solar radiation
Materials		
Building materials:	0.85-0.95	0.65-0.77
Red brick, tile, concrete	0.85-0.95	0.50-0.70
Yellow and buff brick	0.90-0.95	0.50-0.70
Plaster	0.90-0.95	0.82-0.88
Asphalt pavement		
Roofing materials:	0.90-0.95	0.86-0.90
Asphalt	0.25-0.35	0.87-0.91
Galvanized iron, dirty	0.20-0.30	0.64-0.68
Galvanized iron, new	0.85-0.95	0.85-0.90
Roofing paper	0.90-0.98	0.85-0.95
Slate		
Paints:	0.25-0.65	0.30-0.50
Aluminum (bright), gilt	0.95-0.98	0.85-0.95
Black, flat	0.75-0.95	0.65-0.83
Dark (red, brown, green, etc.)	0.95-0.99	0.12-0.25
White, flat		
Metals:	0.02-0.10	0.10-0.40
Aluminum, nickel, chromium (polished)	0.02-0.15	0.30-0.50
Copper, brass, monel metal (polished)	0.20-0.50	0.40-0.65
Dull aluminum, brass, copper, and polished iron	0.60-0.65	0.70-0.80
Iron oxide		





Examples : Standard meridian for USA

75 °W Long. (EST), Eastern standard meridian

90 ° W Long.(CST), Central standard meridian

105 ° W Long. (MST), mountain standard meridian

120 ° W Long. (PST), Pacific standard meridian

Jordan's latitude and longitude

- For Jordan: 31.0167° N, 36.6167° E
- **Jordan's latitude and longitude** is $31^{\circ} 00'$ N and $36^{\circ} 00'$ E



